Quasi-Sasakian 3-manifolds endowed with an η -Einstein Metrics

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Abstract: The aim of this research note is to discuss the characteristics of a 3-dimensional quasi-Sasakian manifold in terms of η -Einstein Solitons. We prove that an η -Einstein soliton on 3-dimensional quasi-Sasakian manifold is an η -Einstein manifold. Moreover, we consider η -Einstein solitons in a 3-dimensional quasi-Sasakian manifold with a Ricci tensor of Codazzi type and cyclic parallel Ricci tensor. Besides these, we discuss, conformally flat and φ -Ricci symmetric η -Einstein soliton in a 3-dimensional quasi-Sasakian manifold. Also, η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold with the curvature condition Q.R = 0 and $Q.W_2 = 0$ have been discussed. Moreover, we furnish an example of η -Einstein solitons in a 3-dimensional quasi-Sasakian-manifold. Finally, we explore an application of η -Einstein solitons in a complete and compact 3-dimensional quasi-Sasakian-manifold to Number theory in terms of homotopy group.

1 Introduction

Geometric flows are important tools for understanding the topological and geometric structures in Riemannian geometry. In 1982, Hamilton [16] introduced that the Ricci solitons move under the Ricci flow simply by diffeomorphisms of the initial metric that is they are stationary points of the Ricci flow given by

$$\frac{\partial g}{\partial t} = -2Ric(g). \tag{1.1}$$

A Ricci soliton (g, V, λ) on a Riemannian manifold is defined by

$$\mathcal{L}_V g + 2S + 2\lambda = 0, \tag{1.2}$$

where S is the Ricci tensor, \mathcal{L}_V is the Lie derivative along the vector field V on M and λ is a real scalar. Ricci soliton is said to be shrinking, steady, or expanding according to $\lambda < 0, \lambda = 0$ and $\lambda > 0$, respectively.

If the vector field V is the gradient of a potential function $-\psi$, where ψ is some smooth function $\psi: M \to \mathbb{R}$, then g is called a gradient Ricci soliton and equation (1.2) assumes the form

$$\nabla\nabla\psi = S + \lambda g. \tag{1.3}$$

It is well known that the quantity

$$a(g,\psi) := R + \left|\nabla\psi^2\right| - \psi$$

must be constant on M and it is often called the *auxiliary constant*. When ψ is constant the gradient Ricci soliton is simply an *Einstein manifold*. Thus Ricci solitons are natural extensions of Einstein metrics, an Einstein manifold with a constant potential function is called a trivial gradient Ricci soliton. Gradient Ricci solitons play an important role in Hamiltonian Ricci flow

as they correspond to self-similar solutions, and often arise as singularity models. They are also related to smooth metric measure spaces, since equation (1.3) is equivalent to ∞ -Bakry-Emery Ricci tensor $Ric\psi = 0$. In physics, a smooth metric space $(M, g, e^{\psi}, dvol)$ with $Ric\psi = \lambda g$ is called quasi-Einstein manifold. Therefore it is important to study the geometry and topology of gradient Ricci solitons and their classifications.

In general, one cannot expect potential function ψ to grow or decay linearly along all directions at infinity, because of the product property: the product of any two gradient steady Ricci solitons is also a gradient steady Ricci soliton. Consider for example (R, g, ψ) , where g is the standard Euclidean metric, $\psi(x_1, x_2) = x_1$. ψ is constant along x_2 direction, so without additional conditions, ψ may not have linear growth at infinity.

In 2016, Catino and Mazzieri introduced the notion of Einstein solitons [8], which generate self-similar solutions to Einstein flow

$$\frac{\partial g}{\partial t} = -2\left(S - \frac{scal}{2}g\right). \tag{1.4}$$

Our interest in studying this equation from different points of view arises from concrete physical problems. On the other hand, gradient vector fields play a central role in Morse-Smale theory. In what follows, after characterizing the manifold of constant scalar curvature via the existence of η -Einstein solitons. The existence of an η -Einstein soliton implies that the manifold is quasi-Einstein. Remark that quasi-Einstein manifolds arose during the study of exact solutions of Einstein field equations.

Sharma [25] initiated the study of Ricci solitons in contact Riemannian geometry. After that, many geometers extensively studied Ricci soliton in almost contact metric manifolds (see [13, 32]). In 2009, Cho and Kimura [9] introduced the notion of η -Ricci solitons and gave a classification of real hypersurfaces in non-flat complex space forms admitting η -Ricci solitons. In [4], Blaga studied the notion of η -Einstein solitons. Siddiqi also studied some properties of η -Einstein solitons in [26, 27, 28, 29, 30, 31] which is closely related to this study.

On the other side, in 1967, Blair [5] introduced the study of quasi-Sasakian structures to unify Sasakian and cosymplectic structures. Quasi-Sasakian manifold can be viewed as an odd dimensional counter part of Kahler structure. In [33] Tanno also studied some facts about the quasi-Sasakian structurs. The properties of quasi-Sasakian manifolds have been studied by several authors, viz., Gonzalez and Chinea [15], Kanemaki [17] and Oubina [19]. In 1990, Kim [18] studied quasi-Sasakian manifolds and proved that fibred Riemannian spaces with invariant fibers normal to the structure vector field do not admit nearly Sasakian or contact structure but a quasi-Sasakian or cosymplectic structure.

Recently, quasi-Sasakian manifolds have been the subject of growing interest in view of finding significant applications to physics, in particular to supergravity and magnetic theory (see [1], [2]). Quasi-Sasakian structures have wide applications in the mathematical analysis of string theory (see[3], [14]). Motivated by the roles of curvature tensor and Ricci tensor of quasi-Sasakian manifolds in string theory [3]. On a 3-dimensional quasi-Sasakian manifold, the structure function β was defined by Olszak [20] and with the help of this function. He has obtained the necessary and sufficient conditions for the manifold to be conformally flat [21]. Next, he has proved that if the manifold is additionally conformally flat with $\beta = constant$, then

(a) the manifold is locally a product of R and a two-dimensional Kächlerian space of constant Gauss curvature (the cosymplectic case), or,

(b) the manifold is of constant positive curvature (the non-cosymplectic case, here the quasi-Sasakian structure is homothetic to a Sasakian structure)[33].

Therefore inspired by the above research literature in the present framework authors extensively studies the η -Einstein solitons on 3-dimensional quasi-Sasakian manifold and a Number theoretic approach to the same setting.

2 Preliminaries

Let M be a (2n + 1)-dimensional connected differentiable manifold endowed with an almost contact metric structure (φ, ξ, η, g) , where φ, ξ, η are tensor field on M of types (1, 1), (1, 0),

(0,1) respectively, such that (see [6], [7])

$$\varphi^2 X = -X + \eta(X)\xi, \qquad (2.1)$$

$$\eta(\xi) = 1, \tag{2.2}$$

$$\eta(X) = g(X,\xi), \quad \varphi\xi = 0, \quad \eta \circ \varphi = 0, \tag{2.3}$$

$$g(\varphi X, \varphi Y) = g(X, Y) - \eta(X)\eta(Y), \quad X, Y \in T(M),$$
(2.4)

where T(M) is the Lie algebra of the vector fields of manifold M.

Let Φ be the fundamental 2-form of M defined by

$$\Phi(X,Y) = g(X,\varphi Y), \quad X,Y \in T(M).$$
(2.5)

Then $\Phi(X,\xi) = 0$, $X \in T(M)$. *M* is said to be quasi-Sasakian if the almost contact structure (φ, ξ, η) is normal and the fundamental 2-form Φ is closed, that is, for every $X, Y \in \Gamma(M)^{2n+1}$, where $\Gamma(M)^{2n+1}$ denotes the module of vector fields on *M*,

$$[\varphi,\varphi](X,Y) + d\eta(X,Y)\xi = 0, \qquad (2.6)$$

$$d\Phi = 0, \quad \Phi(X, Y) = g(X, \varphi Y). \tag{2.7}$$

This was first introduced by Blair [5]. There are many types of quasi-Sasakian structures ranging from the cosymplectic case, $d\eta = 0(rank\eta = 1)$, to the Sasakian case, $\eta(d\eta)n \neq 0(rank\eta = 2n + 1, \varphi = d\eta)$. The 1 form η has rank r = 2p if $(d\eta)^p \neq 0$ and $\eta(d\eta)^p = 0$, and has rank r = 2p + 1 if $(d\eta)^p = 0$ and $\eta(d\eta)^p \neq 0$. We also say that r is the rank of the quasi-Sasakian structure. Blair [5] also proved that there are no quasi-Sasakian structure of even rank. In order to study the properties of quasi-Sasakian manifolds Blair [5] proved some theorems regarding Kählerian manifolds and existence of quasi-Sasakian manifolds.

An almost contact metric manifold M is a 3-dimensional quasi-Sasakian manifold, if and only if its satisfies the following relations [6]

$$\nabla_X \xi = -\beta \varphi X,\tag{2.8}$$

for a certain function β on M, such that $\xi\beta = 0$, ∇ being the operator of the covariant differentiation with respect to the Levi-Civita connection of M. Clearly, such a quasi-Sasakian manifold is cosymplectic if and only if $\beta = 0$.

As a consequence of (2.8), we have [20]

$$(\nabla_X \varphi) Y = \beta[g(X, Y)\xi - \eta(Y)X]$$
(2.9)

$$(\nabla_X \eta) Y = -\beta g(\varphi X, Y) \tag{2.10}$$

for all $X, Y \in T(M)$.

In a 3-dimensional Riemannian manifold, the curvature tensor is given by:

$$R(X,Y)Z = [S(Y,Z)X - S(X,Z)Y + g(Y,Z)QX - g(X,Z)QY]$$
(2.11)

$$-\frac{r}{2}[g(Y,Z)X - g(X,Z)Y],$$
(2.12)

where r is the scalar curvature of M.

Throughout this paper, we consider β as a constant. Let 3-dimensional quasi-Sasakian manifold. Since β is constant The Ricci tensor S and the Ricci operator Q is given by [20]

$$S(X,Y) = \left(\frac{r}{2} - \beta^2\right)g(X,Y) + \left(3\beta^2 - \frac{r}{2}\right)\eta(X)\eta(Y),$$
(2.13)

$$QX = \left(\frac{r}{2} - \beta^2\right)X + \left(3\beta^2 - \frac{r}{2}\right)\eta(X)\xi$$
(2.14)

Moreover, the curvature tensor R, the Ricci tensor S and the Ricci operator Q in a 3-dimensional quasi-Sasakian manifold M also satisfy the following relations [20]

$$R(X,Y)\xi = \beta^2[\eta(Y)X - \eta(X)Y]$$
(2.15)

$$R(\xi, X)Y = \beta^{2}[g(X, Y) - \eta(X)Y]$$
(2.16)

$$S(X,\xi) = 2\beta^2 \eta(X) \tag{2.17}$$

$$Q\xi = 2\beta^2 \xi, \tag{2.18}$$

where g(QX, Y) = S(X, Y). Also, from (2.13) we have

$$S(\varphi X, \varphi Y) = S(X, Y) - 2\beta^2 \eta(X) \eta(Y).$$
(2.19)

3 η -Einstein solitons on $(M, \varphi, \xi, \eta, g)$

Let $(M, \varphi, \xi, \eta, g, \beta)$ be an almost contact metric manifold. Consider the equation

$$\mathcal{L}_{\xi}g + 2S + (2\lambda - scal) + 2\mu\eta \otimes \eta = 0, \tag{3.1}$$

where \mathcal{L}_{ξ} is the Lie derivative operator along the vector field ξ , S is the Ricci curvature tensor field of the metric g, scal is the scalar curvature of the Riemannian metric g and λ and μ are real constants. For $\mu \neq 0$, the data (g, ξ, λ, μ) will be called η -Einstein soliton [4].

Remark that if the scalar curvature *scal* of the manifold is constant, then the η -Einstein soliton $(g, \xi, \lambda - \frac{scal}{2}, \mu)$ reduces to an η -Ricci soliton [9] and, moreover, if $\mu = 0$, to a Ricci soliton [25] $(g, \xi, \lambda - \frac{scal}{2})$. Therefore, the two concepts of η -Einstein soliton [4] and η -Ricci soliton are distinct on manifolds of non constant scalar curvature.

Writing $\mathcal{L}_{\xi}g$ in terms of the Levi-Civita connection ∇ , we obtain [4]:

$$2S(X,Y) = -g(\nabla_X \xi, Y) - g(X, \nabla_Y \xi) - (2\lambda - scal)g(X,Y) - 2\mu\eta(X)\eta(Y),$$
(3.2)

for any $X, Y \in \chi(M)$.

The data (g, ξ, λ, μ) which satisfy the equation (3.1) is said to be an η -Einstein soliton on M [10]. In particular if $\mu = 0$ then (g, ξ, λ) is called *Ricci soliton* [32] and it is called *shrinking*, *steady* or *expanding*, according as λ is negative, zero or positive respectively [9].

Here is an example of η -Einstein soliton on 3-dimensional quasi-Sasakian manifold.

Example 3.1. Let $M = \{(x, y, z) \in \mathbb{R}^3 : z \neq 0\}$ where (x, y, z) are the standard coordinates of \mathbb{R}^3 .

The vector fields are

$$e_1 = \frac{\partial}{\partial z} - y \frac{\partial}{\partial x}, \qquad e_2 = \frac{\partial}{\partial y}, \qquad e_3 = 2 \frac{\partial}{\partial x}$$

Let g be the Riemannian metric defined by

$$g(e_1, e_1) = g(e_2, e_2) = g(e_3, e_3) = 1,$$
 $g(e_1, e_3) = g(e_2, e_3) = g(e_1, e_2) = 0$

that is, the form of the metric becomes Let η be the 1-form defined by $\eta(Z) = g(Z, e_3)$ for any $Z \in \chi(M)$.

Also, let φ be the (1, 1) tensor field defined by

$$\varphi(e_1) = -e_2, \quad \varphi(e_2) = e_1, \quad \varphi(e_3) = 0.$$

Thus using the linearity of φ and g, we have

$$\eta(e_3) = 0, \quad , \eta(e_1) = 0, \quad \eta(e_2) = 0,$$

 $[e_1, e_2] = \frac{1}{2}e_3, \quad [e_2, e_3] = 0, \quad [e_1, e_3] = 0,$
 $\varphi^2 Z = -Z + \eta(Z)e_3$

$$g(\varphi Z, \varphi W) = g(Z, W) - \eta(Z)\eta(W)$$

for any $Z, W \in \chi(M)$.

Then for $e_3 = \xi$, the structure φ, ξ, η, g) defines an almost contact metric structure on M. Let ∇ be the Levi-Civita connection with respect to the metric g, then we have

$$2g(\nabla_X Y, Z) = Xg(Y, Z) + Yg(Z, X) - Zg(X, Y) - g(X, [Y, Z]) -g(Y, [X, Z]) + g(Z, [X, Y]),$$

which is known as Koszul's formula. Using Koszul's formula we have

$$\nabla_{e_1} e_1 = 0, \quad \nabla_{e_1} e_2 = -\frac{1}{4} e_3, \quad \nabla_{e_1} e_3 = \frac{1}{4} e_3,$$
$$\nabla_{e_2} e_1 = \frac{1}{4} e_3, \quad \nabla_{e_2} e_2 = 0, \quad \nabla_{e_2} e_3 = -\frac{1}{4} e_1,$$
$$\nabla_{e_3} e_1 = \frac{1}{4} e_2, \quad \nabla_{e_3} e_2 = -\frac{1}{4} e_1, \quad \nabla_{e_3} e_0 = 0.$$
(3.3)

From (3.3) we find that the structure (φ, ξ, η, g) satisfies the formula (2.8) for $\beta = \frac{1}{4}$ and $\xi = e_3$. Hence the manifold is a 3-dimensional quasi-Sasakian manifold with the constant structure function $\beta = \frac{1}{4}$ [13].

Then the Riemannian and Ricci curvature tensor fields are given by:

$$R(e_1, e_2)e_3 = 0, \quad R(e_2, e_3)e_3 = \frac{1}{16}e_2, \quad R(e_1, e_3)e_3 = \frac{1}{16}e_1,$$

$$R(e_1, e_2)e_2 = -\frac{3}{16}e_1, \quad R(e_2, e_3)e_2 = -\frac{1}{16}e_3, \quad R(e_1, e_3)e_2 = 0,$$

$$R(e_1, e_2)e_1 = \frac{3}{16}e_2, \quad R(e_2, e_3)e_1 = 0, \quad R(e_1, e_3)e = -\frac{1}{16}e_3.$$

From the above expressions of the curvature tensor we obtain

$$S(e_1, e_1) = g(R(e_1, e_2)e_2, e_1) + g(R(e_1, e_3)e_3, e_1) = -\frac{1}{8}$$

similarly we have

$$S(e_1, e_1) = S(e_2, e_2) = -\frac{1}{8}$$
, and $S(e_3, e_3) = \frac{1}{8}$

. In case of η -Einstein soliton, from the relation (3.9) it is sufficient to verify that

$$S(e_i, e_i) = -\left(\lambda - \frac{scal}{2}\right)g(e_i, e_i) - \mu\eta(e_i)\eta(e_i)$$
(3.4)

for all i = 1, 2, 3, and $scal = -\frac{1}{8}$, we get

$$S(e_1, e_1) = -\left(\lambda - \frac{scal}{2}\right)g(e_1, e_1)$$

,

which implies

$$S(e_1, e_1) = -\left(\lambda - \frac{scal}{2}\right) \Rightarrow \lambda = \frac{1}{16}.$$

Also.

$$S(e_3, e_3) = -\left(\lambda - \frac{scal}{2}\right)g(e_3, e_3) - \mu\eta(e_3)\eta(e_3)$$
(3.5)

By using $\lambda = \frac{1}{16}$ and $scal = -\frac{1}{8}$ in (3.5) we obtain $\mu = -\frac{1}{4}$. Therefore, the data $(g, \xi, \lambda - \frac{scal}{2}, \mu)$ is an η -Einstein soliton in 3-dimensional quasi-Sasakian manifold.

For this example we have $\lambda = \frac{1}{16}$, i.e. $\lambda > 0$ so that the η -Einstein soliton in 3-dimensional quasi-Sasakian manifold is expanding.

Proposition 3.2. An η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold is an η -Einstien manifold.

Proof. Let us consider that a 3-dimensional quasi-Sasakian manifold admits a proper η -Einstein soliton $(g, \xi, \lambda - \frac{scal}{2}, \mu)$. Then from the relation (3.1) yields

$$(\mathcal{L}_{\xi}g)(X,Y) + 2S(X,Y) + (2\lambda - scal)g(X,Y) + 2\mu\eta(X)\eta(Y) = 0.$$
(3.6)

It follows that

$$2S(X,Y) = -(\mathcal{L}_{\xi}g)(X,Y) - (2\lambda - scal)g(X,Y) - 2\mu\eta(X)\eta(Y)$$
(3.7)

for all smooth vector fileds X, Y(M). Since ξ is Killing and its integral curves are geodesics in 3-dimensional quasi-Sasakian manifold [20], that is

$$(\mathcal{L}_{\xi}g)(X,Y) = 0. \tag{3.8}$$

Now, using (3.8), we obtain

$$S(X,Y) = -\left(\lambda - \frac{scal}{2}\right)g(X,Y) - \mu\eta(X)\eta(Y).$$
(3.9)

Therefore, we conclude that η -Einstein soliton $(g, \xi, \lambda - \frac{scal}{2}, \mu)$ on a 3-dimensional quasi-Sasakian manifold is an η -Einstein manifold. This complete the proof.

Proposition 3.3. In a 3-dimensional quasi-Sasakian manifold for an η -Einstein soliton we have $\lambda + \mu = 2\beta^2 - \frac{r}{2}.$

Proof. The Ricci tensor of a 3-dimensional quasi-Sasakian manifold is given by

$$S(X,Y) = \left(\frac{r}{2} - \beta^2\right)g(X,Y) + \left(3\beta^2 - \frac{r}{2}\right)\eta(X)\eta(Y),$$
(3.10)

where r is the scalr curvature. Comparing the equation (3.10) with equation (3.9), we obtain $-(\lambda - \frac{scal}{2}) = \frac{1}{2}(r - 2\beta^2)$ and $\mu = -3\beta^2 + \frac{r}{2}$. Since here scal = r is the scalar curvature. From which it follows that $\lambda + \mu = 2\beta^2 - \frac{r}{2}$. This complete the proof.

4 η -Einstein Soliton on 3-dimensional quasi-Sasakian manifolds with Ricci tensor of Coddazi type

In this section, we consider proper η -Einstein soliton on 3-dimensional quasi-Sasakian manifolds with Ricci tensor of Coddazi type. Therefore taking the covariant differentiation of (3.8)with respect to Z we obtain

$$(\nabla_Z S)(X,Y) = -\mu[(\nabla_X \eta)(X)\eta(Y) + \eta(X)(\nabla_Z \eta)(Y)].$$
(4.1)

Using (2.10), we get

$$(\nabla_Z S)(X,Y) = \beta \mu [g(\varphi Z,X)\eta(Y) + \eta(X)g(\varphi Z,Y)].$$
(4.2)

Now, using the fact that the Ricci tensor S of Coddazi type. Then

$$(\nabla_Z S)(X,Y) = (\nabla_Y S)(Z,X). \tag{4.3}$$

Using equations (4.2) and (4.3), we have

$$\beta\mu[g(\varphi Z, X)\eta(Y) + g(\varphi Z, Y)\eta(X)] = \beta\mu[g(\varphi Y, Z)\eta(X) + g(\varphi Y, X)\eta(Z)].$$
(4.4)

Putting $Z = \xi$ in (4.4), we get $\beta \mu = 0$, which is a contradiction. Hence a 3-dimensional quasi-Sasakian manifold with Ricci tensor of Coddazi type does not admit proper η -Einstein soliton. Thus we conclude the following:

Theorem 4.1. A 3-dimensional quasi-Sasakian manifold with Ricci tensor of Coddazi type does not admit a proper η -Einstein soliton.

5 η -Einstein soliton on 3-dimensional quasi-Sasakian manifold with cyclic parallel Ricci tensor

This section deals with a proper η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold with cyclic parallel Ricci tensor. Therefore, we have

$$(\nabla_X S)(Y, Z) + (\nabla_Y S)(Z, X) + (\nabla_Z S)(X, Y) = 0,$$
(5.1)

for all smooth vector fields $X, Y, Z \in \Gamma(TM)$. Using (3.9) in (5.1), we have

$$\beta\mu[g(Y,\varphi X)\eta(Z) + g(Z,\varphi X)\eta(Y) + g(Z,\varphi Y)\eta(X)$$
(5.2)

$$+g(X,\varphi Y)\eta(Z) + g(X,\varphi Z)\eta(Y) + g(Y,\varphi Z)\eta(X)] = 0.$$

Putting $X = \xi$ in (refs6), we get

$$\beta \mu g(\varphi Y, Z) = 0. \tag{5.3}$$

It follow that $\beta \mu = 0$, since $(\beta \neq 0)$ or $\mu = 0$, which is a contradiction. Thus we can state the following theorem:

Theorem 5.1. A 3-dimensional quasi-Sasakian manifold with cyclic parallel Ricci tensor does not admit a proper η -Eisntein soliton.

6 φ -Ricci Symmetric η -Einstein soliton on 3-dimensional quasi-Sasakian manifold

In this section, we study φ -Ricci symmetric proper η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold. A quasi-Sasakian manifold is said to be φ -Ricci symmetric [13] if

$$\varphi^2(\nabla_X Q)Y = 0 \tag{6.1}$$

for all smooth vector fields X, Y on M.

The Ricci tensor for an η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold is given by

$$S(X,Y) = -\left(\lambda - \frac{scal}{2}\right)g(X,Y) - \mu\eta(X)\eta(Y).$$
(6.2)

Then it follows that

$$QX = -\left(\lambda - \frac{scal}{2}\right)X - \mu\eta(X)\xi \tag{6.3}$$

for all smooth vector filed x on M. Now, taking covariant derivative of (s10), we obtain

$$(\nabla_X Q)Y = \beta \mu g(\varphi X, Y) + \beta \mu \eta(Y)\varphi X.$$
(6.4)

Applying φ^2 on both side of the equation (6.4), we get

$$\varphi^2(\nabla_X Q)Y = \beta \mu \eta(Y)\varphi^2 X. \tag{6.5}$$

From (6.1) and (6.5) we have

$$\beta \mu = 0. \tag{6.6}$$

Since $\beta \neq 0$. Therefore $\mu = 0$, which is a contradiction. Therefore a φ -Ricci symmetric 3-dimensional quasi-Sasakian manifold does not admit a proper η -Einstein soliton. Thus we can state the following theorem:

Theorem 6.1. A φ -Ricci symmetric 3-dimensional quasi-Sasakian manifold does not admit a proper η -Einstein soliton.

7 Conformally flat η -Einstein soliton on 3-dimensional quasi-Sasakian manifold

In this section we study conformally flat proper η -Einstein soliton in a 3-dimensional quasi-Sasakian manifold. Therefore [?]

$$(\nabla_X S)(Y,Z) - (\nabla_Y S)(X,Z) = \frac{1}{4} [g(Y,Z)dr(X) - g(X,Z)dr(Y)].$$
(7.1)

Using (3.9) in (7.1)

$$\beta\mu[g(Y,\varphi X)\eta(Z) + g(Z,\varphi X)\eta(Y)?g(X,\varphi Y)\eta(Z)?g(Z,\varphi Y)\eta(X)]$$
(7.2)

$$= \frac{1}{4} [g(Y,Z)dr(X) - g(X,Z)dr(Y)].$$

Putting $X = \xi$ in (7.2), we get

$$4\beta\mu g(\varphi Y, Z) = -\eta(Z)dr. \tag{7.3}$$

It follow that $4\beta\mu\varphi Y = -dr(Y)\xi$. This implies that $\beta\mu\varphi^2 Y = 0$. Hence $\beta\mu = 0$. Since $\beta \neq 0$ or $\mu = 0$, which is a contradiction. Therefore a conformally flat 3-dimensional quasi-Sasakian manifold does not admit a proper η -Einstein soliton. Thus we can state the following theorem:

Theorem 7.1. A conformally flat 3-dimensional quasi-Sasakian manifold does not admit a proper η -Einstein soliton.

8 η -Einstein soliton on 3-dimensional quasi-Sasakian manifold satisfying the curvature condition Q.R = 0

In this section, we study the proper η -Einstein soliton in a 3-dimensional quasi-Sasakian manifold satisfying the curvature condition Q.R = 0. Therefore

$$(Q.R)(X,Y)Z = 0 (8.1)$$

for all smooth vector fileds X, Y, Z on M. The explicit form of the above equation is

$$Q(R(X,Y)Z) - R(QX,Y)Z - R(X,QY)Z - R(X,Y)QZ = 0.$$
(8.2)

Using (3.9) in (8.2) we have

$$-\lambda_1 R(X,Y)Z - \mu\eta(R(X,Y)Z)\xi + \lambda_1 R(X,Y)Z + \mu\eta(X)R(\xi,Y)Z$$
(8.3)

$$\begin{aligned} &+\lambda_{1}R(X,Y)Z + \mu\eta(Y)R(X,\xi)Z + \lambda_{1}R(X,Y)Z + \mu\eta(Z)R(X,Y)\xi = 0, \\ \text{where } \lambda_{1} = \left(\lambda - \frac{scal}{2}\right). \text{ Using (2.11) in (8.3) we have} \\ &-\mu g(Y,Z)\eta(QX)\xi + \mu g(X,Z)\eta(QY)\xi - \mu\eta(X)\eta(QZ)Y \\ &+mu\eta(X)g(Y,Z)Q\xi - 2\mu\eta(X)\eta(Z)QY + r\mu\eta(X)\eta(Z)Y \\ &+\mu\eta(Y)\eta(QZ)X + 2\mu\eta(Y)\eta(Z)QX - \mu\eta(Y)g(X,Z)Q\xi \\ &-r\mu\eta(Y)\eta(Z)X + \mu\eta(Z)\eta(QY)X - \mu\eta(Z)\eta(QX)Y \\ &+2\lambda_{1}[S(Y,Z)X - S(X,Z)Y + g(Y,Z)QX - g(X,Z)QY] \\ &-\frac{r}{2}[g(Y,Z)X - g(X,Z)Y] = 0. \end{aligned}$$

Again using
$$(3.9)$$
 in (8.4) we obtain

$$\lambda_{1}\mu g(Y,Z)\eta(X)\xi + \mu^{2}g(Y,Z)\eta(X)\xi - \mu g(Y,Z)\eta(-\lambda_{1}X - \mu\eta(X)\xi)\xi$$

$$+\mu g(X,Z)\eta(-\lambda_{1}X - \mu\eta(Y)\xi)\xi - \eta(X)\eta(-\lambda_{1}Z - \mu\eta(Z)\xi)Y$$

$$+\mu\eta(X)g(Y,Z)(-\lambda_{1}\xi - \mu\xi) - 2\mu\eta(X)\eta(Z)(\lambda_{1}Y - \mu\eta(Y)\xi + r\mu\eta(X)\eta(Z)Y$$

$$+\mu\eta(Y)\eta(-\lambda_{1}\xi - \mu\xi) - r\mu\eta(Y)\eta(Z)X + \mu\eta(Z)\eta(-\lambda_{1}Y - \mu\eta(Y)\xi)X$$

$$\mu\eta(Z)\eta(-\lambda_{1}X - \mu\eta(X)\xi)Y + 2\lambda_{1}[(-\lambda_{1}g(Y,Z) - \mu\eta(Y)\eta(Z))X - (-\lambda_{1}g(X,Z)$$

$$-\mu\eta(X)\eta(Z))Y + g(Y,Z)(-\lambda_{1}X - \mu\eta(X)\xi) - g(X,Z)(-\lambda_{1}Y - \mu\eta(Y)\xi)$$

$$-\frac{r}{2}(g(Y,Z)X - g(X,Z)Y)] = 0.$$
(8.5)

from which it follows that

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$$4\lambda_1\mu\eta(X)\eta(Z)Y + 2\mu^2\eta(X)\eta(Z)Y - 4\lambda_1\mu\eta(Y)\eta(Z)X + r\mu\eta(X)\eta(Z)Y$$
(8.6)

$$-r\mu\eta(Y)\eta(Z)X + 2\lambda_{1}[(-\lambda_{1}g(Y,Z) - \mu\eta(X)\eta(Z))X - (-\lambda_{1}g(X,Z) - \mu\eta(X)\eta(Z))Y + g(Y,Z)(-\lambda_{1}X - \mu\eta(X)\xi) - g(X,Z)(-\lambda_{1}Y - \mu\eta(Y)\xi) - \frac{r}{2}(g(Y,Z)X - g(X,Z)Y)]$$

$$-2\mu^2\eta(Y)\eta(Z)X + \lambda_1\mu g(Y,Z)\eta(X)\xi + \mu^2 g(Y,Z)\eta(X)\xi = 0.$$

Putting $X = Z = \xi$ in (8.6) we get

$$4\lambda_{1}\mu Y - 4\lambda_{1}\mu\eta(Y)\xi + r\mu Y - r\mu\eta(Y)\xi$$

$$+2\lambda_{1}[-\lambda_{1}\eta(Y)\xi - \mu\eta(Y)\xi + \lambda_{1}Y + \mu Y - \lambda_{1}\eta(Y)\xi$$

$$(8.7)$$

$$-\mu\eta(Y)\xi + \lambda_1 Y + \mu\eta(Y)\xi - \frac{1}{2}\eta(Y)\xi + \frac{1}{2}Y$$

$$-2\mu^2\eta(Y)\xi + \lambda_1\mu\eta(Y)\xi + \mu^2\eta(Y)\xi = 0.$$

Taking inner product of (8.7) with ξ we get

$$\mu(\lambda_1 + \mu)\eta(Y) = 0. \tag{8.8}$$

By hypothesis $\mu \neq 0$, therefore from the above equation we get

$$\lambda_1 + \mu = 0, ie.$$
 $\lambda - \frac{scal}{2} + \mu = 0$

Moreover, for an η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold, we have

$$\lambda + \mu = 2\beta^2 - \frac{r}{2}$$

Thus in this case $\lambda = \beta^2 - \frac{r}{4}$ and $\mu = \beta^2 - \frac{r}{4}$. Therefore we have the following:

Theorem 8.1. An proper η -Einstein soliton on 3-dimensional quasi-Sasakain manifold Sasakian satisfying the curvature condition Q.R = 0 is of the type $(g, V, \frac{\beta^2}{2}, \frac{\beta^2}{2})$.

9 η -Einstein soliton on 3-dimensional quasi-Sasakian manifold satisfying the curvature condition $Q.W_2 = 0$.

 W_2 -curvature tensor filed is the curvature tensor introduced by Pokhariyal and Mishra [24].

$$W_2(X,Y)Z = R(X,Y)Z + \frac{1}{dimM - 1}[g(X,Z)QY - g(Y,Z)QX].$$
(9.1)

In this section, we study proper η -Einstein soliton on a 3-dimensional quasi-Sasakian manifold satisfying the curvature condition $Q.W_2 = 0$. Therefore

$$(Q.W_2)(X,Y)Z = 0. (9.2)$$

for all smooth vector filed X, Y, Z on M. The explicit form of the equation (9.2) is

$$Q(W_2(X,Y)Z) - W_2(QX,Y)Z - W_2(X,QY)Z - W_2(X,Y)QZ = 0.$$
(9.3)

Using (3.9) in (9.3) we have

$$-\lambda_1 W_2(X,Y)Z - \mu\eta(W_2(X,Y)Z)\xi + \lambda_1 W_2(X,Y)Z + \mu\eta(X)W_2(\xi,Y)Z$$
(9.4)

$$+\lambda_1 W_2(X,Y)Z + \mu \eta(Y) W_2(X,\xi)Z + \lambda_1 W_2(X,Y)Z + \mu \eta(Z) W_2(X,Y)\xi = 0,$$

where $\lambda_1 = (\lambda - \frac{scal}{2})$. Using (9.1) in (9.4) we have

$$-\mu g(Y,Z)\eta(QX)\xi + \mu g(X,Z)\eta(QY)\xi - \mu \eta(X)\eta(QZ)Y$$
(9.5)

$$-\frac{\mu}{2}g(X,Z)\eta(QY)\xi + \frac{\mu}{2}g(Y,Z)\eta(QX)\xi$$

 $+\mu\eta(X)g(Y,Z)Q\xi - 2\mu\eta(X)\eta(Z)QY + r\mu\eta(X)\eta(Z)Y$

$$\frac{1}{2}\mu\eta(X)\eta(Z)QY - \frac{1}{2}\mu\eta(X)g(Y,Z)Q\xi$$

$$+\mu\eta(Y)\eta(QZ)X + 2\mu\eta(Y)\eta(Z)QX - \mu\eta(Y)g(X,Z)Q\xi$$
$$\frac{\mu}{2}\eta(Y)g(X,Z)Q\xi - \frac{\mu}{2}\eta(Y)\eta(Z)QX$$
$$-r\mu\eta(Y)\eta(Z)X + \mu\eta(Z)\eta(QY)X - \mu\eta(Z)\eta(QX)Y$$
$$+\frac{\mu}{2}\eta(Z)\eta(Y)QY - \frac{\mu}{2}\eta(Z)\eta(Y)QX$$
$$+2\lambda_{1}[S(Y,Z)X - S(X,Z)Y + g(Y,Z)QX - g(X,Z)QY]$$
$$-\frac{r}{2}[g(Y,Z)X - g(X,Z)Y] + \frac{1}{2}[g(X,Z)QY - g(Y,Z)QX] = 0.$$

Again using (3.9) in (9.5) we obtain

$$\frac{\lambda_{1}\mu}{2}g(Y,Z)\eta(X)\xi + \frac{\mu^{2}}{2}g(Y,Z)\eta(X)\xi - \frac{\mu}{2}g(Y,Z)\eta(-\lambda_{1}X - \mu\eta(X)\xi)\xi \qquad (9.6)$$

$$+\frac{\mu}{2}g(X,Z)\eta(-\lambda_{1}X - \mu\eta(Y)\xi)\xi - \eta(X)\eta(-\lambda_{1}Z - \mu\eta(Z)\xi)Y$$

$$+\mu\eta(X)g(Y,Z)(-\lambda_{1}\xi - \mu\xi) - \frac{5}{2}\mu\eta(X)\eta(Z)(\lambda_{1}Y - \mu\eta(Y)\xi + r\mu\eta(X)\eta(Z)Y$$

$$+\mu\eta(Y)\eta(-\lambda_{1}\xi - \mu\xi) - r\mu\eta(Y)\eta(Z)X + \mu\eta(Z)\eta(-\lambda_{1}Y - \mu\eta(Y)\xi)X$$

$$-\mu\eta(Z)\eta(-\lambda_{1}X - \mu\eta(X)\xi)Y + 2\lambda_{1}[(-\lambda_{1}g(Y,Z) - \mu\eta(Y)\eta(Z))X - (-\lambda_{1}g(X,Z))$$

$$-\mu\eta(X)\eta(Z))Y + g(Y,Z)(-\lambda_{1}X - \mu\eta(X)\xi) - g(X,Z)(-\lambda_{1}Y - \mu\eta(Y)\xi)$$

 $-\frac{r}{2}(g(Y,Z)X - g(X,Z)Y)] + \frac{1}{2}[[g(X,Z)(-\lambda_1Y - \mu\eta(Y)\xi - g(Y,Z)(-\lambda_1X - \mu\eta(X)\xi]] = 0.$ from which it follows that

$$\frac{9}{2}\lambda_{1}\mu\eta(X)\eta(Z)Y + 2\mu^{2}\eta(X)\eta(Z)Y - 3\lambda_{1}\mu\eta(Y)\eta(Z)X + r\mu\eta(X)\eta(Z)Y$$
(9.7)
$$-r\mu\eta(Y)\eta(Z)X + 2\lambda_{1}[(-\lambda_{1}g(Y,Z) - \mu\eta(X)\eta(Z))X - (-\lambda_{1}g(X,Z) - \mu\eta(X)\eta(Z))Y + g(Y,Z)(-\lambda_{1}X - \mu\eta(X)\xi) - g(X,Z)(-\lambda_{1}Y - \mu\eta(Y)\xi) - \frac{r}{2}(g(Y,Z)X - g(X,z)Y)]$$
$$-\lambda_{1}^{2}g(X,Z)Y + \lambda_{1}\mu\eta(Y)g(X,Z)\xi + \lambda_{1}^{2}g(Y,Z)X + \lambda_{1}\mu\eta(X)g(Y,Z)\xi$$
$$-2\mu^{2}\eta(Y)\eta(Z)X + \frac{3}{2}\mu^{2}\eta(X)\eta(Y)\eta(Z)\xi = 0.$$

Putting $X = Z = \xi$ in (9.7) we get

$$\frac{9}{2}\lambda_{1}\mu Y + 2\mu^{2}Y - 3\lambda - 1\mu\eta(Y)\xi + r\mu Y - r\mu\eta(Y)\xi$$
(9.8)

$$+2\lambda_1[-\lambda_1\eta(Y)\xi - \mu\eta(Y)\xi + \lambda_1Y + \mu Y - \lambda_1\eta(Y)\xi$$

$$-\mu\eta(Y)\xi + \lambda_1 Y + \mu\eta(Y)\xi - \frac{1}{2}\eta(Y)\xi + \frac{1}{2}Y - \lambda_1^2\eta(Y)\xi + \lambda_1^2 Y$$
$$-2\mu^2\eta(Y)\xi + \frac{3}{2}\mu^2\eta(Y)\xi = 0.$$

Taking inner product of (9.8) with ξ we get

$$\mu\left(\frac{5}{2}\lambda_1 + \frac{3}{2}\mu\right)\eta(Y) = 0.$$
(9.9)

By hypothesis $\mu \neq 0$, therefore from the above equation we get

$$\frac{5}{2}\lambda_1 + \frac{3}{2}\mu = 0, ie.$$
 $\frac{1}{2}(5\lambda - scal) + \frac{3}{2}\mu = 0$

Moreover for an η -Einstein soliton on 3-dimensional quasi-Sasakian manifold we have

$$\lambda + \mu = 2\beta^2 - \frac{r}{2}$$

Thus in this case $\lambda = 2\beta^2$ and $\mu = -2\beta^2$. Therefore we have the following:

Theorem 9.1. An proper η -Einstein soliton on 3-dimensional quasi-Sasakain manifold Sasakian satisfying the curvature condition $Q.W_2 = 0$ is of the type $(g, V, 2\beta^2, -2\beta^2)$.

10 Application of η -Einstein soliton on quasi-Sasakain manifold to Algebraical Number Theory

In 2008, Wylie [34] proved that if (M, g) is a complete Riemannian manifold satisfying (1.2), then M has a finite fundamental group (for more information see Theorem 1.1 [34]).

Remark 10.1. According to Theorem 1 and Theorem 2, proved by Rustanov (see [23]), he has shown that with a non-zero scalar curvature a quasi-Sasakian manifold is complete, it is compact and has a finite fundamental group.

Remark 10.2. [22] A finite fundamental group is the first and simplest homotopy group. The fundamental group is homotopy invariant, topological spaces are homotopy equivalent and have isomorphic fundamental groups. The fundamental group or a homotopy group of a topological space X is denoted by $\pi_1(X)$.

In view of (3.1), non-zero scalar curvature *scal*, implies the existence η -Einstein soliton on 3-dimensional quasi-Sasakain manifold. Therefore, in light of (3.1), Remark 10.1, and Remark 10.2 we gain the following outcomes:

Theorem 10.3. If a 3-dimensional quasi-Sasakain manifold (M^3, g) admits an η -Einstein soliton with non-zero scalar curvature scal, then, it is complete, compact and has a finite fundamental group.

Theorem 10.4. If a 3-dimensional quasi-Sasakain manifold (M^3, g) admits an η -Einstein soliton with non-zero scalar curvature scal, then, it is complete, compact and has a homotopy group $\pi_1(M^3)$.

Next, we have the following discussion, which is based on the Number theoretic approach with the homotopy group of a space.

Remark 10.5. For a prime number p, the homotopy p-exponent of a topological space \mathcal{T} , denoted by $Exp_p(U)$, is defined to be a largest $e \in \mathbb{N} = \{0, 1, 2, \dots\}$ such that some homotopy group $\Phi_j(\mathcal{T})$ has an element of order p^e . Cohen et al. [10] proved that the

$$Exp_p(\mathcal{S}^{2n+1}) = n \quad if \quad p \neq 2.$$

For a prime number p and a integer z, the p-adic order of z is given by $Ord_p(z) = sup \{z \in \mathbb{N} : p^z | z\}$.

In 2007, Davis and Sun derived an interesting inequality in terms of a homotopy group (Theorem 1.1 Page 2 [12]). According to them, for any prime p and $z = 2, 3, \cdots$ some homotopy group $\pi_i(\mathcal{T})$ contains an element of order $p^{n-1+Ord_p(\lfloor n/p \rfloor!)}$, i.e., then the strong and elegant lower bound for homotopy p-exponent of homotopy group is

$$Exp_p(\pi_i(\mathcal{T})) \ge n - 1 + Ord_p\left(\left\lfloor \frac{n}{p} \right\rfloor!\right), \tag{10.1}$$

where $\pi_i(\mathcal{T})$ is a homotopy group of degree n.

Therefore, using Davis and Sun result (Theorem 1.1 [12]) with Theorem 10.4, we gain an interesting inequality

Theorem 10.6. For any prime number p and $s = 2, 3, \dots, a$ homotopy group $\pi_1(M^3)$ of a complete 3-dimensional quasi-Sasakain manifold (M^3, g) admits an η -Einstein soliton with non-zero scalar curvature scal, contains an element of order $p^{s-1+Ord_p(\lfloor s/p \rfloor!)}$, we obtain the inequality

$$Exp_p(\pi_1(M^3)) \ge n - 1 + Ord_p\left(\left\lfloor \frac{s}{p} \right\rfloor!\right).$$
(10.2)

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