

Notes on Generalized Quasi Einstein Manifolds

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Abstract The Ricci tensor of a generalized quasi-Einstein manifold M has the form

$R_{ij} = \alpha g_{ij} + \beta A_i A_j + \gamma B_i B_j$ where A and B are two orthonormal forms. It is proved that, among other results, the two generators are closed if and if the Ricci tensor is a Codazzi tensor. M reduces to be quasi-Einstein if M is Ricci symmetric and the second generator is not parallel. M is Einstein if its Ricci tensor is symmetric and the two generators are not parallel. Sufficient conditions on a generalized quasi-Einstein GRW space-time to be either Einstein or nearly quasi-Einstein are given. Finally, a generalized quasi-Einstein pseudo-Ricci symmetric manifold are investigated.

1 Introduction

A pseudo-Riemannian manifold M is called Einstein if its Ricci tensor satisfies

$$R_{ij} = \alpha g_{ij},$$

where α is a constant. It is well-known that $\alpha = \frac{R}{n}$ where R is the scalar curvature of M [3]. A quasi-Einstein manifold was introduced by Chaki and R. K. Maity[5] as a generalization of Einstein manifolds. Such manifolds have a Ricci tensor of the form

$$R_{ij} = \alpha g_{ij} + \beta A_i A_j,$$

where α and β are scalars and A is a unit 1-form. Quasi-Einstein Lorentzian manifolds are called perfect fluid space-times whenever A is time-like[20]. Some basic geometric properties of quasi-Einstein space-time were obtained in[19]. De and Ghosh investigated generalized quasi-Einstein manifold in [7] as a new generalization of Einstein manifolds and an extension of Chaki’s notion. In a generalized quasi-Einstein manifold, the Ricci tensor satisfies

$$R_{ij} = \alpha g_{ij} + \beta A_i A_j + \gamma B_i B_j, \tag{1.1}$$

where α, β, γ are scalars, A and B are two orthonormal 1-forms[2, 8, 12]. The forms A and B are called the generators of generalized quasi-Einstein manifolds. De and Gazi in [9] introduced the notion of a nearly quasi-Einstein manifold in which Ricci curvature tensor is of the form

$$R_{ij} = \alpha g_{ij} + \beta E_{ij},$$

where E_{ij} is $(0, 2)$ symmetric tensor. A pseudo-Ricci symmetric manifold, denoted by $(PRS)_n$, is a pseudo-Riemannian manifold whose Ricci tensor satisfies

$$(\nabla_X R)(Y, Z) = 2A(X)R(Y, Z) + A(Y)R(X, Z) + A(Z)R(Y, X),$$

where A is a non-zero 1-form and ∇ is the Livi-Civita connection[4]. Locally it is,

$$\nabla_k R_{ij} = 2A_k R_{ij} + A_i R_{kj} + A_j R_{ik}.$$

Generalized quasi-Einstein manifolds have been investigated in mathematics and physics literature. These manifolds portray a generalization of Einstein manifolds and an extension of

quasi-Einstein manifolds. Perfect fluid space-times are pictured out as quasi-Einstein manifolds. In [11], the authors studied generalized quasi-Einstein manifolds admitting W_2 -curvature tensor whereas generalized quasi-Einstein manifolds admitting the conharmonic curvature tensor are studied by Prakasha and Venkatesha in [16]. Non-trivial examples of generalized quasi-Einstein manifolds were considered in Section 6 of [11]. Sular et al gave some classification results of generalized quasi-Einstein manifolds[21]. Recently, in [13], the authors proved that a generalized quasi-Einstein pseudo Ricci symmetric manifolds can not admit a Codazzi Ricci tensor and the generators of these manifolds are not parallel.

In the present paper, some geometric properties of generalized quasi-Einstein manifolds are investigated. Second, Results on generalized quasi-Einstein warped product manifold are discussed. Next, it is proved that a generalized quasi-Einstein GRW space-time is either Einstein or nearly quasi-Einstein. Finally, a generalized quasi-Einstein pseudo-Ricci symmetric manifold is considered.

2 On generalized quasi-Einstein manifolds

Proposition 2.1. *The scalar curvature of a generalized quasi-Einstein manifold with generators A and B is given by*

$$R = (n - 2) \alpha + (A^i A^j + B^i B^j) R_{ij}.$$

Proof. Let M be a generalized quasi-Einstein manifold, that is,

$$R_{ij} = \alpha g_{ij} + \beta A_i A_j + \gamma B_i B_j,$$

where $A_i A^i = 1$, $B_i B^i = 1$ and $A_i B^i = 0$. A contraction with g^{ij} implies

$$R = n\alpha + \beta + \gamma, \quad (2.1)$$

which means that M has a constant scalar curvature R . Contracting Equation (1.1) with $A^i A^j$ and then by $B^i B^j$, one gets

$$A^i A^j R_{ij} = \alpha + \beta, \quad (2.2)$$

$$B^i B^j R_{ij} = \alpha + \gamma \quad (2.3)$$

Thus,

$$R = (n - 2) \alpha + (A^i A^j + B^i B^j) R_{ij}.$$

□

Theorem 2.2. *Let M be a generalized quasi-Einstein manifold. Then the generators A and B are closed if the Ricci tensor is a Codazzi tensor. Moreover, the covariant derivative of the Ricci curvature tensor is given by*

$$\nabla_j R_{ik} = 2\beta A_k (\nabla_j A_i) + 2\gamma B_k (\nabla_j B_i).$$

Proof. The covariant derivative of Equation (1.1) implies

$$\nabla_k R_{ij} = \beta (\nabla_k A_i) A_j + \beta A_i (\nabla_k A_j) + \gamma (\nabla_k B_i) B_j + \gamma B_i (\nabla_k B_j).$$

Also,

$$\nabla_i R_{kj} = \beta (\nabla_i A_k) A_j + \beta A_k (\nabla_i A_j) + \gamma (\nabla_i B_k) B_j + \gamma B_k (\nabla_i B_j).$$

Then, the Codazzi deviation tensor D is

$$\begin{aligned} D_{kij} &= \nabla_k R_{ij} - \nabla_i R_{kj} \\ &= \beta (\nabla_k A_i) A_j + \beta A_i (\nabla_k A_j) + \gamma (\nabla_k B_i) B_j + \gamma B_i (\nabla_k B_j) \\ &\quad - [\beta (\nabla_i A_k) A_j + \beta A_k (\nabla_i A_j) + \gamma (\nabla_i B_k) B_j + \gamma B_k (\nabla_i B_j)]. \end{aligned}$$

Two contractions with A^j and B^j give

$$A^j D_{kij} = \beta [\nabla_k A_i - \nabla_i A_k],$$

$$B^j D_{kij} = \gamma [\nabla_k B_i - \nabla_i B_k].$$

If R_{ij} is a Codazzi tensor, then the generators are closed, that is, $\nabla_k A_i = \nabla_i A_k$ and $\nabla_k B_i = \nabla_i B_k$.

Conversely, assume that A and B are closed, then

$$\begin{aligned} 0 &= \beta A_i (\nabla_k A_j) + \gamma B_i (\nabla_k B_j) - [\beta A_k (\nabla_i A_j) + \gamma B_k (\nabla_i B_j)] \\ &= \nabla_j (\beta A_k A_i + \gamma B_k B_i) - 2\beta A_k (\nabla_j A_i) - 2\gamma B_k (\nabla_j B_i) \\ &= \nabla_j R_{ik} - 2\beta A_k (\nabla_j A_i) - 2\gamma B_k (\nabla_j B_i), \end{aligned}$$

Thus the covariant derivative of R_{ik} is

$$\nabla_j R_{ik} = 2\beta A_k (\nabla_j A_i) + 2\gamma B_k (\nabla_j B_i).$$

□

Theorem 2.3. *A Ricci symmetric generalized quasi-Einstein manifold M reduces to be quasi-Einstein if B is not parallel. M reduces to be Einstein if the two generators are not parallel together.*

Proof. The covariant derivative of the Ricci tensor for generalized quasi-Einstein is

$$\beta (\nabla_i A_k) A_j + \beta A_k (\nabla_i A_j) + \gamma (\nabla_i B_k) B_j + \gamma B_k (\nabla_i B_j) = \nabla_i R_{kj}.$$

Assume that the Ricci curvature tensor is symmetric that is, $\nabla_i R_{kj} = 0$. Thus

$$\beta (\nabla_i A_k) A_j + \beta A_k (\nabla_i A_j) + \gamma (\nabla_i B_k) B_j + \gamma B_k (\nabla_i B_j) = 0.$$

Two different contractions with A^k and B^k imply

$$\begin{aligned} \beta (\nabla_i A_j) &= 0, \\ \gamma (\nabla_i B_j) &= 0. \end{aligned}$$

Assume that B is not parallel, then $\gamma = 0$, and hence Equation (1.1) leads to

$$R_{ij} = \alpha g_{ij} + \beta A_i A_j,$$

which means that M is quasi-Einstein. Moreover, if A and B are not parallel, then $\gamma = \beta = 0$ and hence Equation (1.1) reduces to

$$R_{ij} = \alpha g_{ij}.$$

Thus, M is Einstein. □

3 Generalized quasi-Einstein warped product manifold

Let (\bar{M}, \bar{g}) and (\tilde{M}, \tilde{g}) be two pseudo-Riemannian manifolds with dimensions $\dim \bar{M} = \bar{n}$, $\dim \tilde{M} = \tilde{n}$ and let $F : \bar{M} \rightarrow (0, \infty)$ be a smooth positive function on \bar{M} . Consider the product manifold $\bar{M} \times \tilde{M}$ with its natural projections $\pi : \bar{M} \times \tilde{M} \rightarrow \bar{M}$ and $\eta : \bar{M} \times \tilde{M} \rightarrow \tilde{M}$. Then the warped product manifold $M = \bar{M} \times_F \tilde{M}$ is the manifold $\bar{M} \times \tilde{M}$ furnished with metric

$$g = \bar{g} \oplus f^2 \tilde{g}.$$

The manifold \bar{M} is called the base manifold of M whereas \tilde{M} is called the fiber manifold of M [6].

Now, Let $a, b, c, d, \dots \in \{1, \dots, \bar{n}\}$ denote the basis vector fields on a neighborhood \bar{U} of the base manifold \bar{M} whereas $\alpha, \beta, \gamma, \delta, \dots \in \{\bar{n} + 1, \dots, n\}$ denote the basis vector fields on a neighborhood \tilde{U} of the fiber manifold \tilde{M} . Likewise, $i, j, k, l, \dots \in \{1, \dots, n\}$ denote the basis vector fields on a neighborhood $\bar{U} \times \tilde{U}$ of the warped product manifold. The local components of the metric $g = \bar{g} \times_F \tilde{g}$, are $g_{ab} = \bar{g}_{ab}$, $g_{\alpha j} = 0$ and $g_{\alpha\beta} = F \tilde{g}_{\alpha\beta}$. The local components Γ_{ij}^h of the Livi-Civita connection on the warped product $M = \bar{M} \times_F \tilde{M}$ are as follows

$$\begin{aligned} \Gamma_{bc}^a &= \bar{\Gamma}_{bc}^a, & \Gamma_{\beta\gamma}^\alpha &= \tilde{\Gamma}_{\beta\gamma}^\alpha, \\ \Gamma_{\alpha\beta}^a &= \frac{1}{2} \bar{g}^{ab} F_b \tilde{g}_{\alpha\beta}, & \Gamma_{a\beta}^\alpha &= \frac{1}{2F} F_a \delta_\beta^\alpha, \\ \Gamma_{ab}^\alpha &= \Gamma_{\alpha b}^a = 0 \end{aligned}$$

where $F_a = \partial_a F = \frac{\partial F}{\partial x^a}$.

The local components of the Riemannian curvature tensor of the warped product $M = \bar{M} \times_F \tilde{M}$ are given as

$$\begin{aligned} R_{abcd} &= \bar{R}_{abcd}, \\ R_{\alpha\beta\gamma\delta} &= F\bar{R}_{\alpha\beta\gamma\delta} - \frac{1}{4}P\tilde{G}_{\alpha\beta\gamma\delta}, \\ R_{\alpha a\beta b} &= \frac{-1}{2}T_{ab}\tilde{g}_{\alpha\beta}, \end{aligned}$$

where $P = \bar{g}^{ab}F_aF_b$, and $\tilde{G}_{\alpha\beta\gamma\delta} = \tilde{g}_{\alpha\gamma}\tilde{g}_{\beta\delta} - \tilde{g}_{\alpha\delta}\tilde{g}_{\beta\gamma}$. T is a $(0, 2)$ tensor and its local components are

$$T_{ab} = \bar{\nabla}_b F_a - \frac{1}{2F}F_aF_b.$$

Locally, the Ricci curvature R_{ij} of the warped product $M = \bar{M} \times_F \tilde{M}$ has the following components

$$R_{ab} = \bar{R}_{ab} - \frac{\tilde{n}}{2F}T_{ab}, \quad (3.1)$$

$$R_{\alpha\beta} = \bar{R}_{\alpha\beta} - \left[\text{tr}(T) + \frac{\tilde{n}-1}{2F}\bar{\Delta}F \right] \tilde{g}_{\alpha\beta}, \quad (3.2)$$

where $\text{tr}(T) = \bar{g}^{ab}T_{ab}$.

Theorem 3.1. *Let $M = \bar{M} \times_F \tilde{M}$ be a generalized quasi-Einstein warped product manifold, then the base manifold \bar{M} is generalized quasi-Einstein.*

Proof. Let M be a warped product manifold. Then

$$R_{ab} = \bar{R}_{ab} - \frac{\tilde{n}}{2F}T_{ab}.$$

Assume that M is $G(QE)_n$, thus

$$R_{ab} = \alpha\bar{g}_{ab} + \beta\bar{A}_a\bar{A}_b + \gamma\bar{B}_a\bar{B}_b.$$

The last two Equations imply

$$\bar{R}_{ab} - \frac{\tilde{n}}{2F}T_{ab} = \alpha\bar{g}_{ab} + \beta\bar{A}_a\bar{A}_b + \gamma\bar{B}_a\bar{B}_b.$$

If $T_{ab} = 0$, then

$$\bar{R}_{ab} = \alpha\bar{g}_{ab} + \beta\bar{A}_a\bar{A}_b + \gamma\bar{B}_a\bar{B}_b,$$

which illustrates that \bar{M} is generalized quasi-Einstein manifold. \square

Theorem 3.2. *If $M = \bar{M} \times_F \tilde{M}$ is a generalized quasi-Einstein warped product manifold, then the fiber manifold \tilde{M} is generalized quasi-Einstein.*

Proof. Let M be a warped product manifold. Thus

$$R_{\alpha\beta} = R_{\alpha\beta} = \bar{R}_{\alpha\beta} - \left[\text{tr}(T) + \frac{\tilde{n}-1}{2F}\bar{\Delta}F \right] \tilde{g}_{\alpha\beta}.$$

Now, assume that M is $G(QE)_n$, then

$$R_{\alpha\beta} = \tilde{g}_{\alpha\beta} + \beta\tilde{A}_\alpha\tilde{A}_\beta + \gamma\tilde{B}_\alpha\tilde{B}_\beta.$$

Using the last two Equations one gets

$$\bar{R}_{\alpha\beta} = \left[\text{tr}(T) + \frac{\tilde{n}-1}{2F}\bar{\Delta}F + \alpha \right] \tilde{g}_{\alpha\beta} + \beta\tilde{A}_\alpha\tilde{A}_\beta + \gamma\tilde{B}_\alpha\tilde{B}_\beta,$$

which shows that \tilde{M} is $G(QE)_n$ warped product manifold. \square

4 Generalized quasi-Einstein GRW space-times

A generalized Robertson-Walker space-time (or a GRW space-time) is the warped product $M = I \times_f M^*$ of an open connected interval $(I, -dt^2)$ with warping function $f : I \rightarrow R^+$. A Lorentzian manifold M is a generalized Robertson-Walker space-time if and only if M possess a unit time-like vector field u_i such that

$$\nabla_k u_i = \varphi (g_{ki} + u_k u_i), \quad (4.1)$$

$$R_{ij} u^j = \xi u_i, \quad (4.2)$$

where φ and ξ are scalar functions. Vector fields satisfying Equation (4.1) are called torse-forming and those obeying Equation (4.2) are eigenvectors of the Ricci tensor with eigenvalue ξ [14, 15].

Theorem 4.1. *The following statements are true in a generalized quasi-Einstein GRW space-time M ,*

- (i) M reduces to be a nearly quasi-Einstein manifold if u^i is not orthogonal to the both generators,
- (ii) M reduces to be Einstein if u^i is orthogonal to both the generators, provided $\varphi \neq 0$ and $\xi = \text{const}$.

Proof. Assume that M is a generalized quasi-Einstein manifold, that is

$$R_{ij} = \alpha g_{ij} + \beta A_i A_j + \gamma B_i B_j.$$

Multiplying by u^i , one gets

$$(\xi - \alpha) u_j = \beta u^i A_i A_j + \gamma u^i B_i B_j$$

The two Contractions of the both sides with A^i and B^i imply

$$(\xi - \alpha - \beta) u^i A_i = 0,$$

$$(\xi - \alpha - \gamma) u^i B_i = 0.$$

There are two different possible cases

- (i) If u^i is not orthogonal to the both generator that is, $u^i A_i = u^i B_i \neq 0$, then $\beta = \gamma$ and hence the Ricci tensor becomes.

$$R_{ij} = \alpha g_{ij} + \beta (A_i A_j + B_i B_j).$$

Since $(A_i A_j + B_i B_j) = E_{ij}$, we have

$$R_{ij} = \alpha g_{ij} + \beta E_{ij},$$

where $(A_i A_j + B_i B_j) = E_{ij}$ is a symmetric tensor.

- (ii) If u^i is orthogonal to the both generators, that is, $u^i A_i = u^i B_i = 0$, and hence

$$u^i R_{ij} = \alpha u_j,$$

The covariant derivative of the Ricci tensor of M is

$$\nabla_k R_{ij} = \beta [A_i (\nabla_k A_j) + A_j (\nabla_k A_i)] + \gamma [B_i (\nabla_k B_j) + B_j (\nabla_k B_i)]$$

contracting with u^i and using the condition $u^i A_i = u^i B_i = 0$, then

$$u^i \nabla_k R_{ij} = 0,$$

$$\nabla_k u^i R_{ij} - R_{ij} \nabla_k u^i = 0$$

If M is generalized Robertson-Walker space-time, then the following two equations hold

$$\nabla_k u^i = \varphi (\delta_k^i + u_k u^i),$$

$$R_{ij} u^i = \xi u_j$$

Thus

$$\begin{aligned} \nabla_k (\xi u_j) - R_{ij} \varphi (\delta_k^i + u_k u^i) &= 0 \\ u_j \nabla_k (\xi) + \xi \nabla_k u_j - \varphi \delta_k^i R_{ij} - \varphi u_k u^i R_{ij} &= 0 \\ \xi \varphi (g_{kj} + u_k u_j) - \varphi R_{kj} - \xi \varphi u_k u_j &= 0 \\ u_j \nabla_k (\xi) + \varphi (\xi g_{kj} - R_{kj}) &= 0. \end{aligned}$$

If $\xi = \text{const}$, then the above Equation becomes

$$\varphi (\xi g_{kj} - R_{kj}) = 0.$$

If $\varphi \neq 0$, then M is an Einstein manifold. □

5 Generalized quasi-Einstein (PRS) $_n$ manifold

Assume that M is generalized quasi-Einstein, then

$$R_{ij} = \alpha g_{ij} + \beta B_i B_j + \gamma D_i D_j, \quad (5.1)$$

By contracting with g^{ij} , one gets

$$R = n\alpha + \beta + \gamma, \quad (5.2)$$

which means that the scalar curvature is constant.

Let M be a pseudo-Ricci symmetric manifold, then the covariant derivative of the Ricci tensor is

$$\nabla_k R_{ij} = 2A_k R_{ij} + A_i R_{kj} + A_j R_{ik}, \quad (5.3)$$

Multiplying the both sides of Equation (5.3) by g^{kj} , one may get the divergence of the Ricci tensor as follows

$$\nabla_k R_i^k = 3A^j R_{ij} + A_i R.$$

It is well-known that

$$\nabla_k R_i^k = \frac{1}{2} \nabla_i R.$$

Thus,

$$\nabla_i R = 6A^j R_{ij} + 2A_i R. \quad (5.4)$$

A different contraction of Equation (5.3) by g^{ij} , it is

$$\nabla_k R = 2A_k R + 2A^j R_{kj}, \quad (5.5)$$

By solving these two equations, one gets

$$A^j R_{kj} = 0, \quad (5.6)$$

$$\nabla_k R = 2A_k R. \quad (5.7)$$

Equations (5.7) and (5.2) imply

$$R = 0.$$

Theorem 5.1. *The scalar curvature of a generalized quasi-Einstein (PRS) $_n$ manifold is zero.*

Assume that the Ricci tensor of a generalized quasi-Einstein $(PRS)_n$ manifold M is cyclic parallel[1, 10], that is,

$$\nabla_j R_{ki} + \nabla_k R_{ji} + \nabla_i R_{kj} = 0.$$

Equation (5.3) implies

$$A_k R_{ij} + A_j R_{ik} + A_i R_{jk} = 0. \quad (5.8)$$

Contracting Equation (5.8) with A^k and use Equation (5.6), one may obtain

$$aR_{ij} = 0, \quad (5.9)$$

where $a = A^k A_k$. Thus

$$R_{ij} = 0,$$

which means that M is Ricci flat. Then we have the following theorem.

Theorem 5.2. *A generalized quasi-Einstein $(PRS)_n$ manifold with cyclic parallel Ricci tensor is Ricci flat.*

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